


*Transformation laws
of
redshift, flux and luminosity
in
generic spacetimes*



Maurício O. Calvão*, B. L. Lago**,
R. R. R. Reis* & B. B. Siffert*

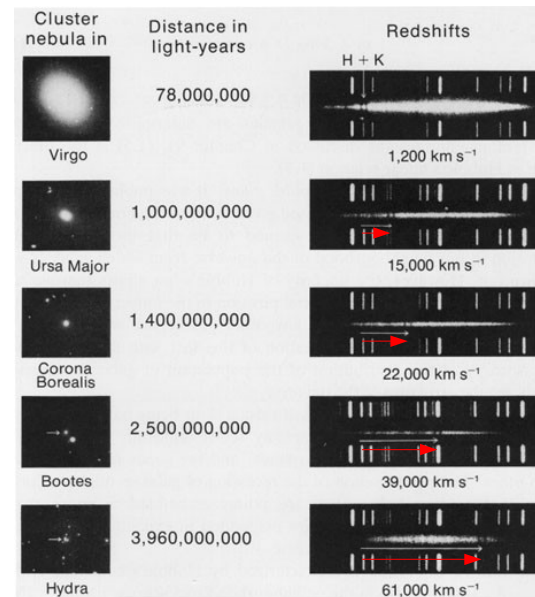
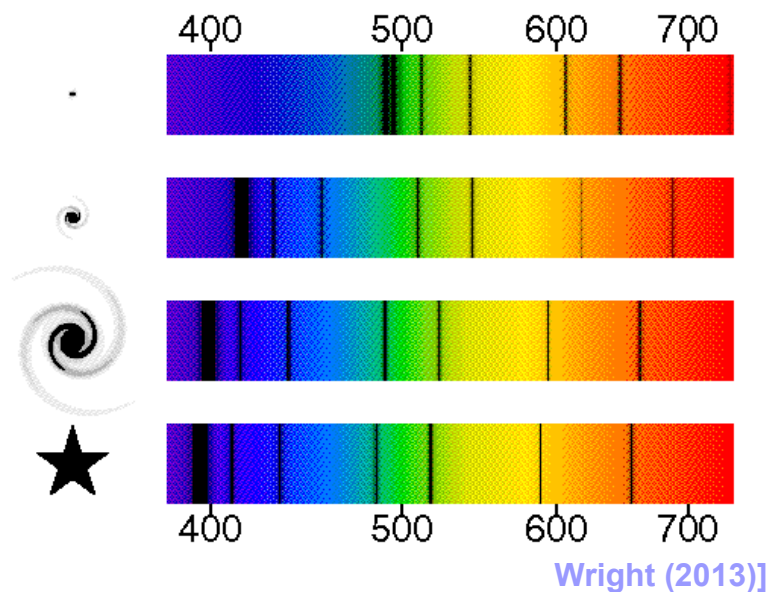
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I. INTRODUCTION

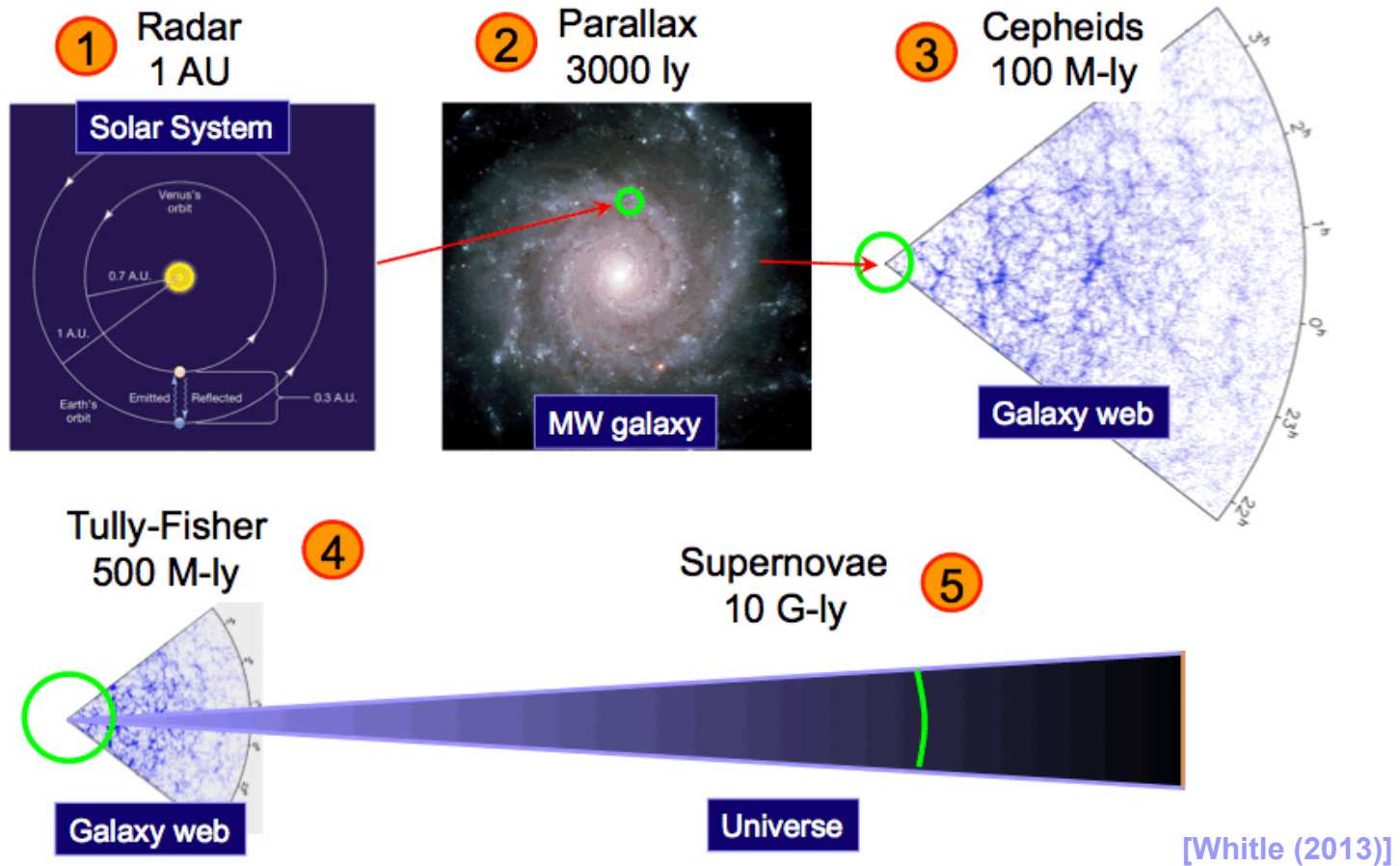
I.1. MOTIVATION

redshift: Doppler, gravitational, cosmological



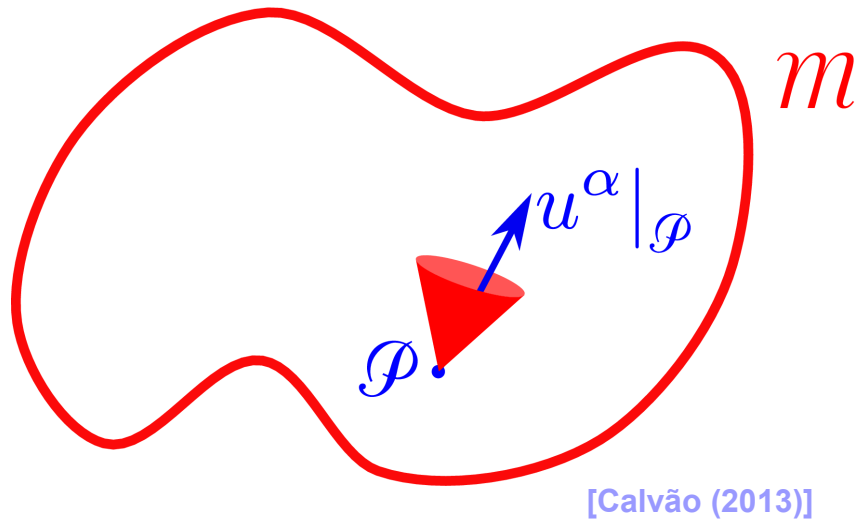
Schombert (2013)]

flux and luminosities



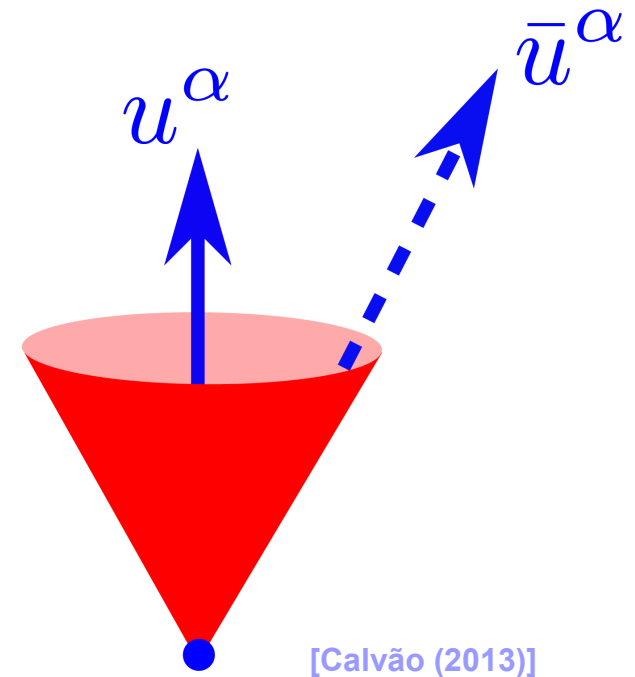
I.2. BASIC CONCEPTS

instantaneous observer



$$h^{\alpha\beta} := g^{\alpha\beta} - u^\alpha u^\beta$$

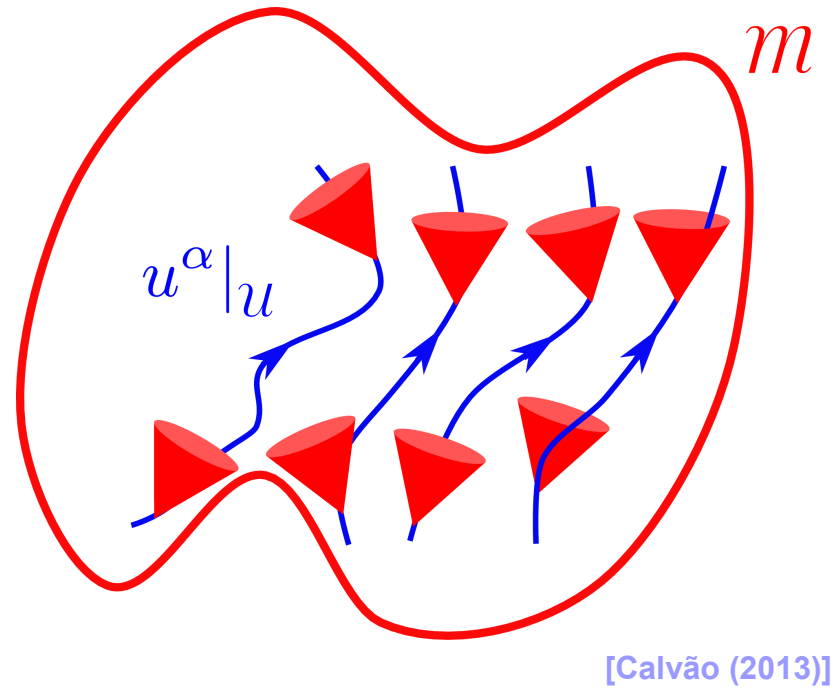
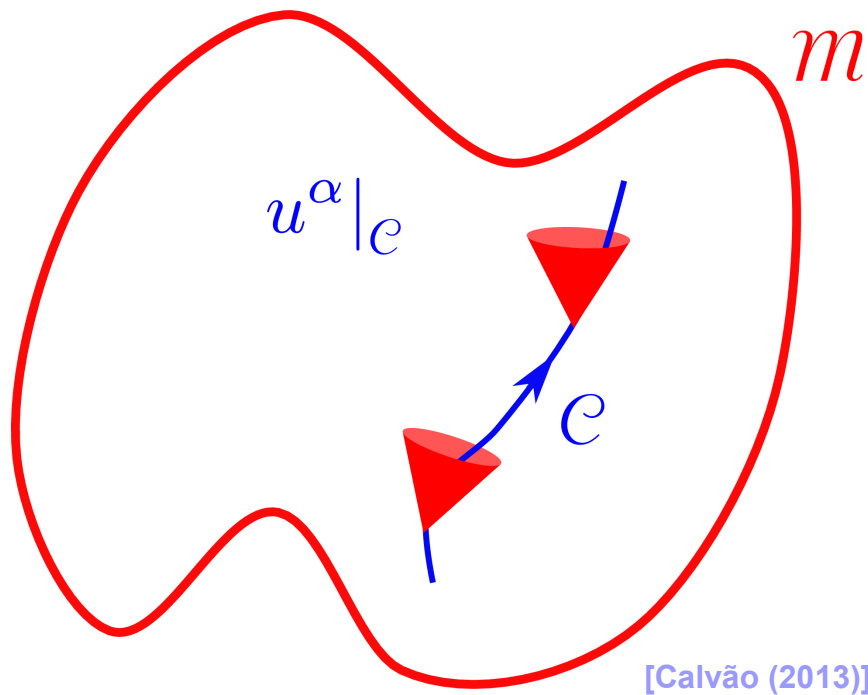
local Lorentz boost



$$\bar{u}^\alpha = \gamma (u^\alpha + v e^\alpha)$$

observer

frame of reference



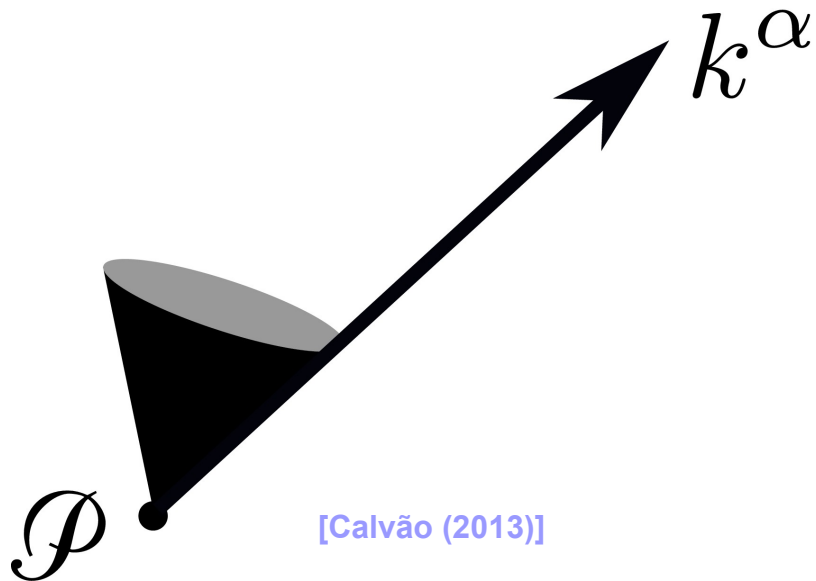
$$a^\alpha := \frac{Du^\alpha}{D\tau}$$

$$u_{\alpha;\beta} = a_\alpha u_\beta + \frac{1}{3}\Theta h_{\alpha\beta} + \sigma_{\alpha\beta} + \omega_{\alpha\beta}$$



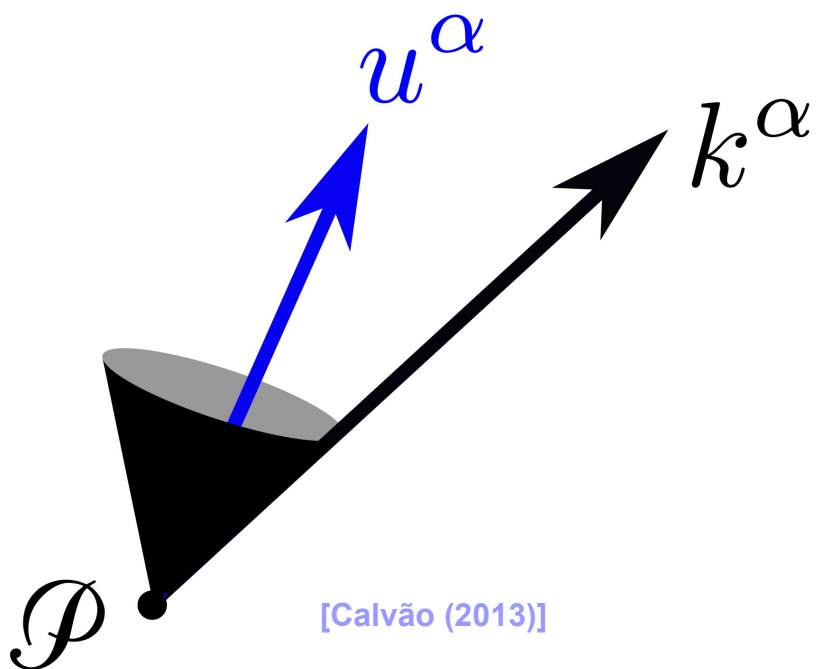
II. REDSHIFT

II.1. LOCAL (OR POINTWISE)



II. REDSHIFT

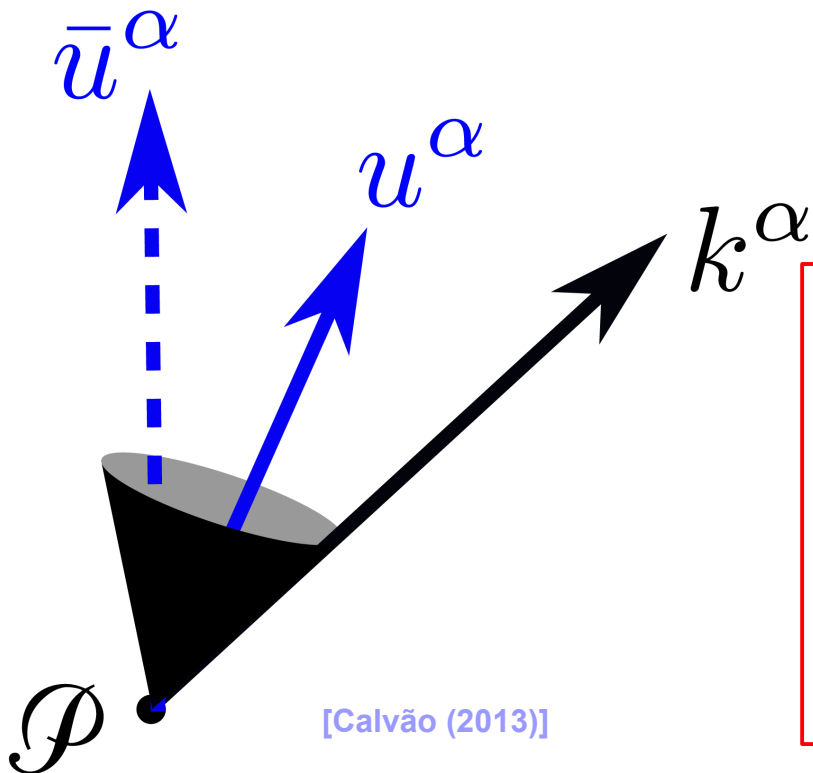
II.1. LOCAL (OR POINTWISE)



$$k^\alpha = E (u^\alpha + n^\alpha)$$

II. REDSHIFT

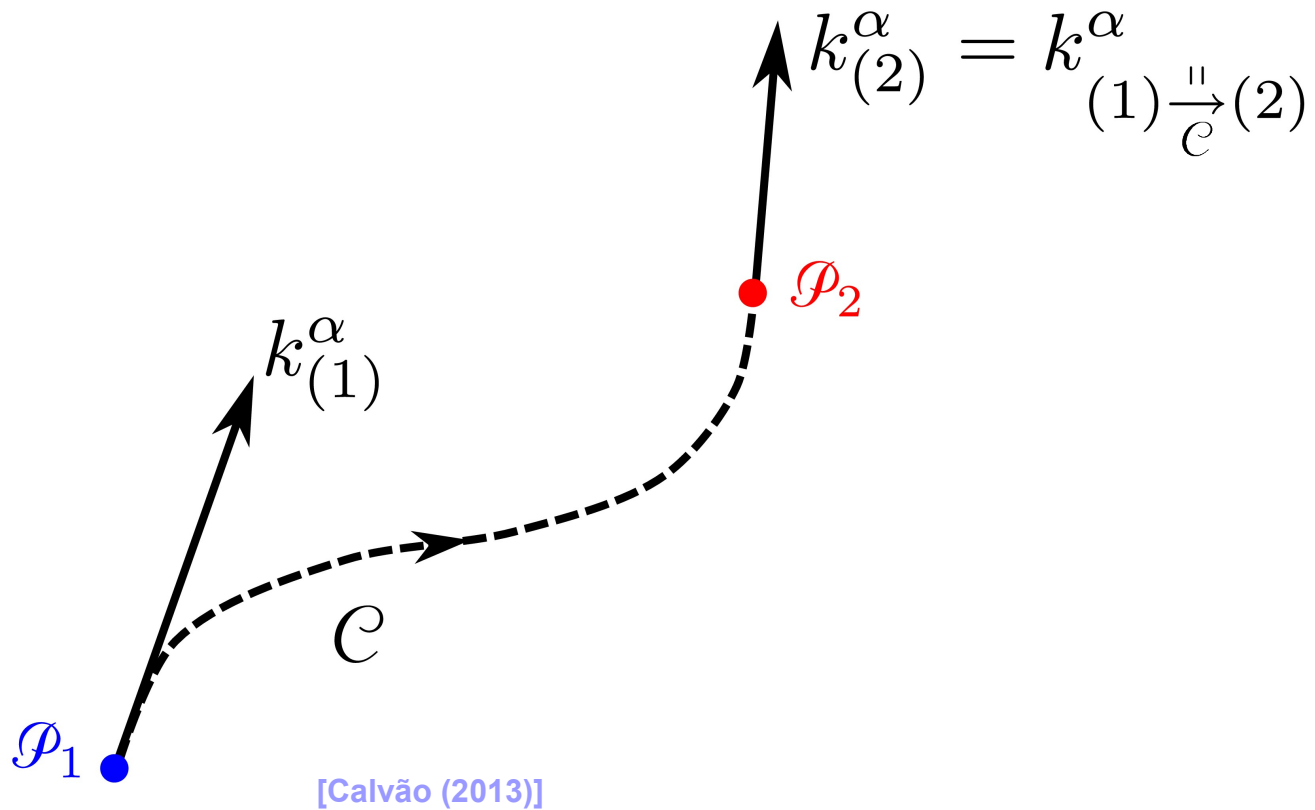
II.1. LOCAL (OR POINTWISE)



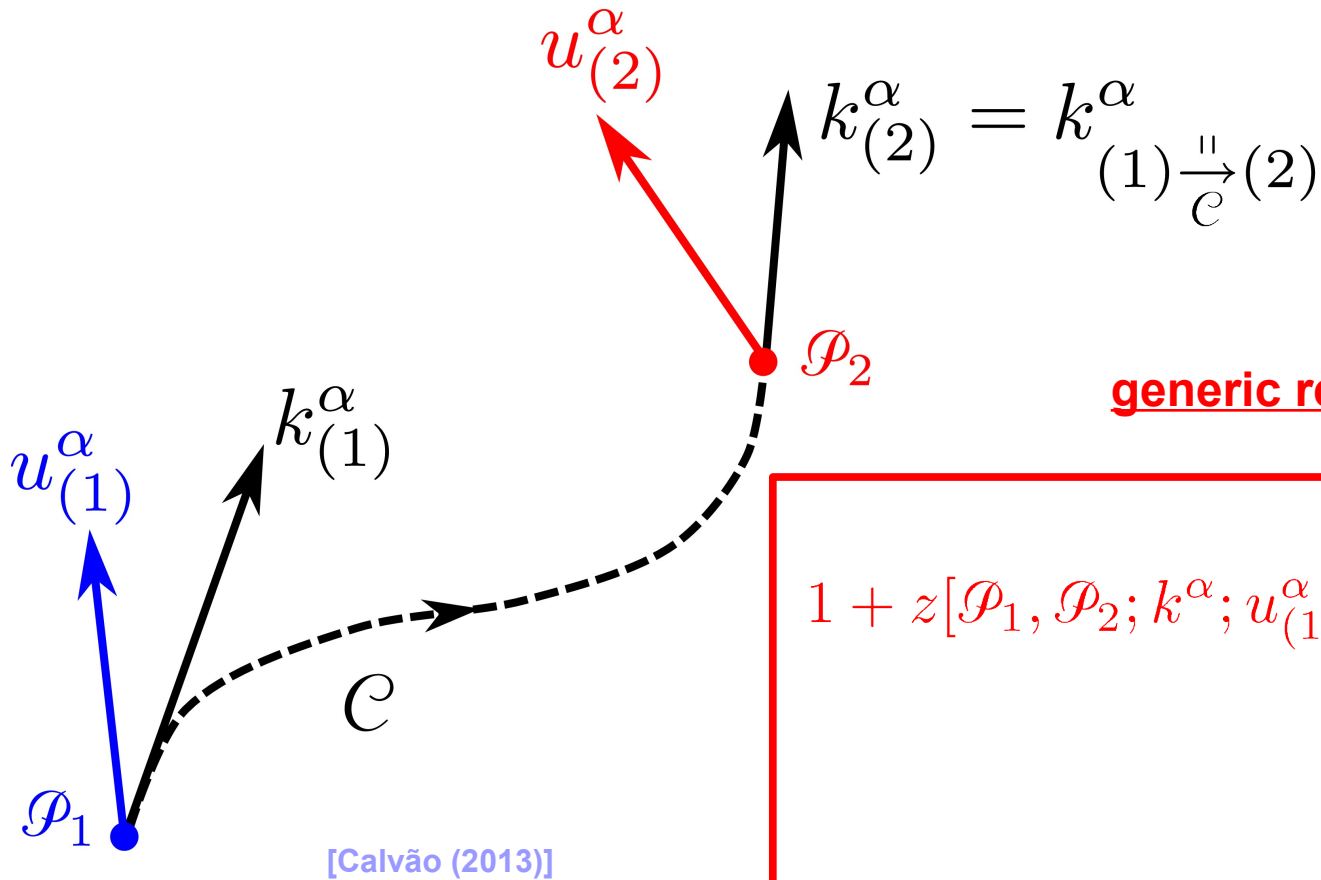
$$\bar{E} = \gamma E (1 + v n^\alpha e_\alpha)$$

$$1 + z[\mathcal{P}; k^\alpha; u^\alpha \rightarrow \bar{u}^\alpha] = \frac{\bar{\lambda}}{\lambda} = \frac{k^\alpha u_\alpha}{k^\beta \bar{u}_\beta} \Big|_{\mathcal{P}}$$

II.2. GLOBAL (OR AT A DISTANCE)



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generic redshift law:

$$\begin{aligned}
 1 + z[\mathcal{P}_1, \mathcal{P}_2; k^\alpha; u_{(1)}^\alpha \rightarrow u_{(2)}^\alpha] &= \frac{\lambda_2}{\lambda_1} \\
 &= \frac{(k^\alpha u_\alpha)|_{\mathcal{P}_1}}{(k^\beta u_\beta)|_{\mathcal{P}_2}}
 \end{aligned}$$

II.2. GLOBAL (OR AT A DISTANCE)

in particular, for any spacetime:

$$\exists u_{(2)}^\alpha \mid u_{(2)}^\alpha = u_{(1)}^\alpha \stackrel{||}{\underset{c}{\rightarrow}} (2) \Rightarrow z = 0$$



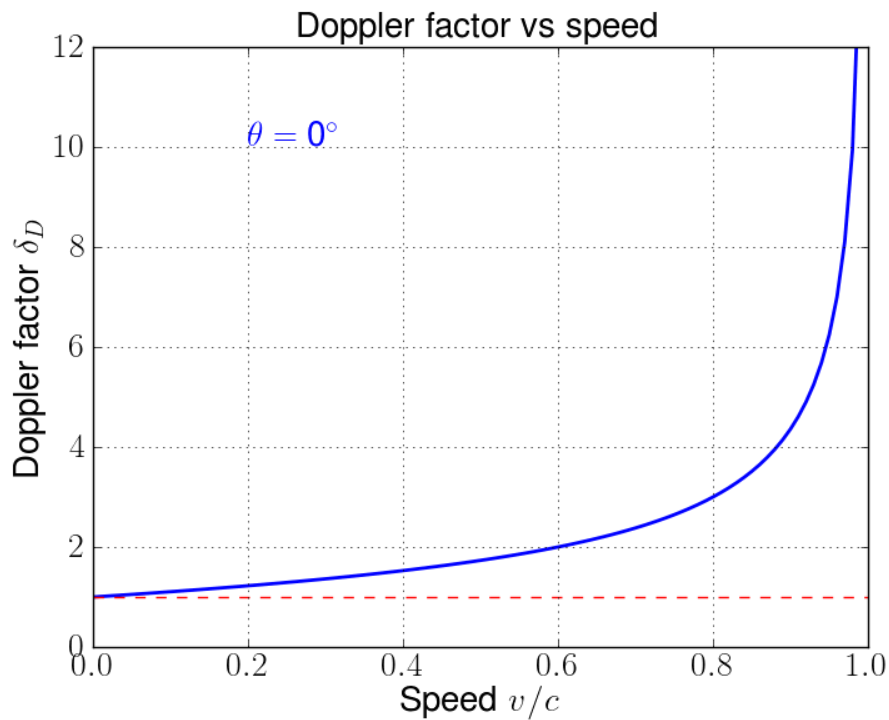
challenge to expanding space paradigm!

basis for consistent definition of absolute magnitude!

II.3. DOPPLER FACTOR

$$\delta_D(v, \theta) := [\gamma(v) (1 - v \cos \theta)]^{-1}$$

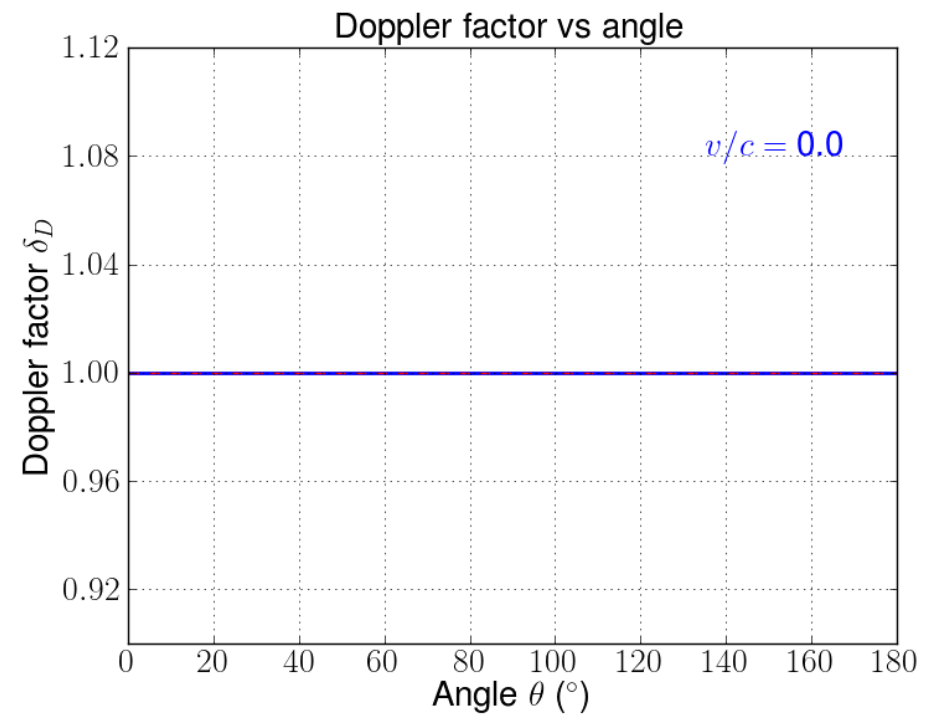
$$\Rightarrow 1 + z(v, \theta) = \delta_D(v, \theta)$$



[Calvão (2013)]

II.3. DOPPLER FACTOR

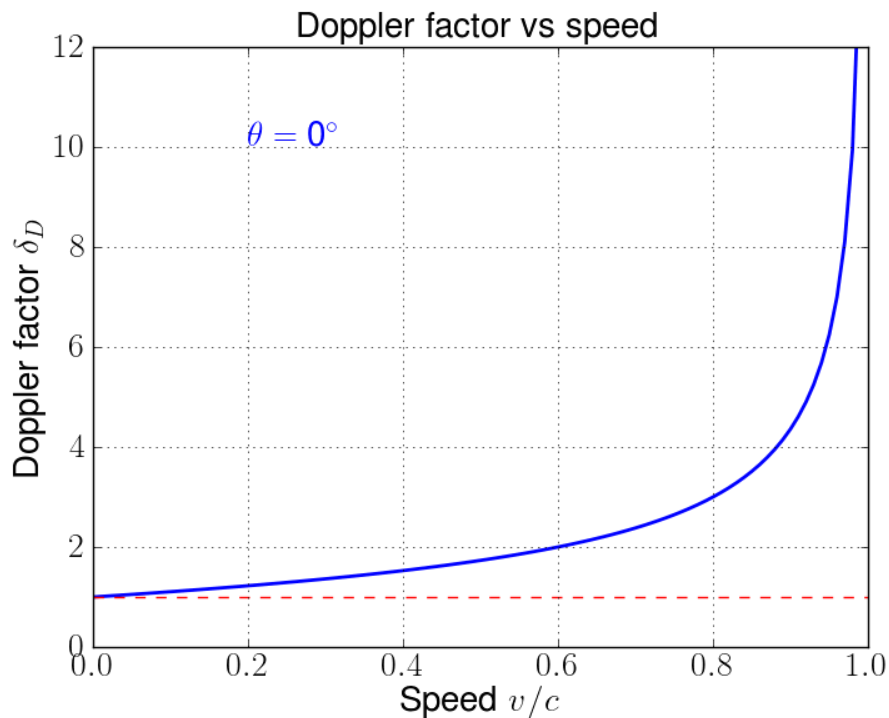
$$\delta_D(v, \theta) := [\gamma(v) (1 - v \cos \theta)]^{-1} \Rightarrow 1 + z(v, \theta) = \delta_D(v, \theta)$$



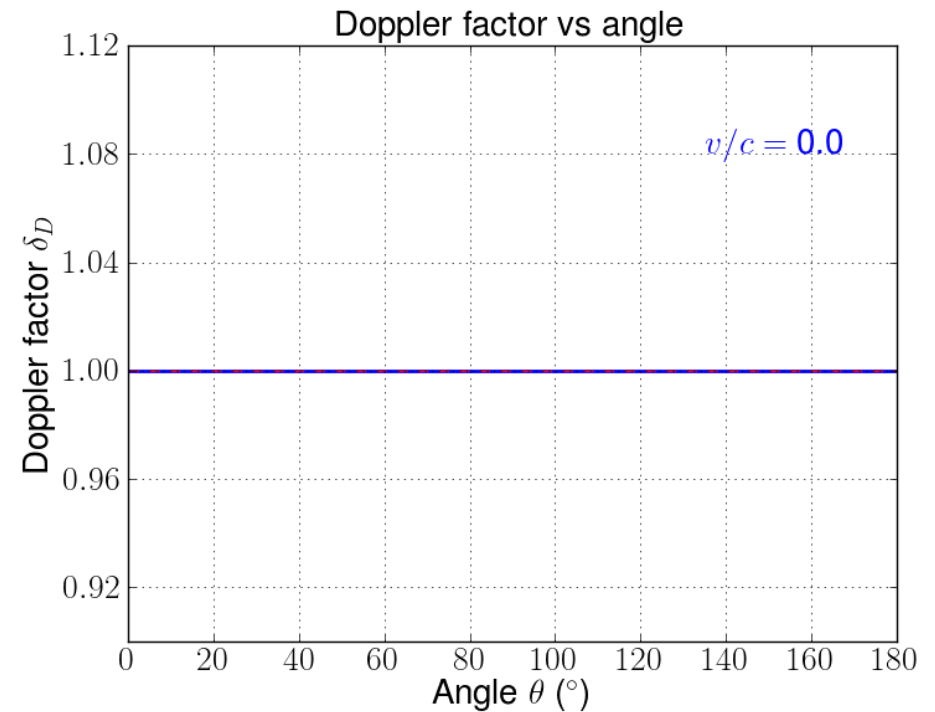
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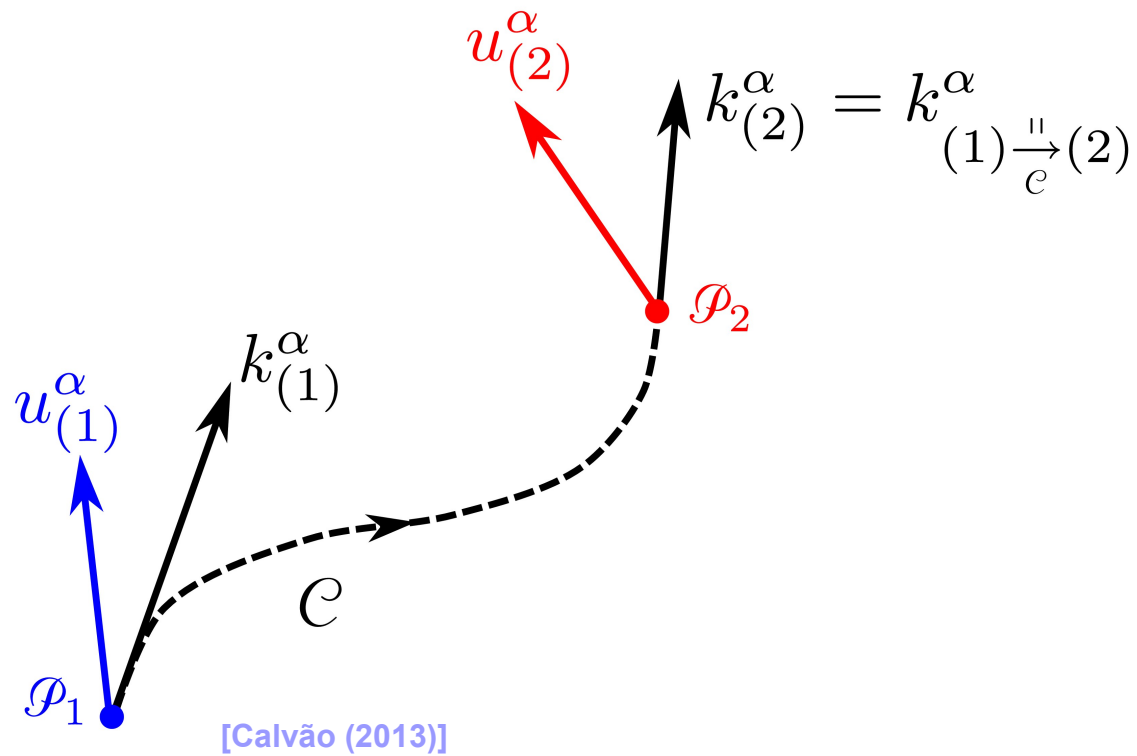


[Calvão (2013)]

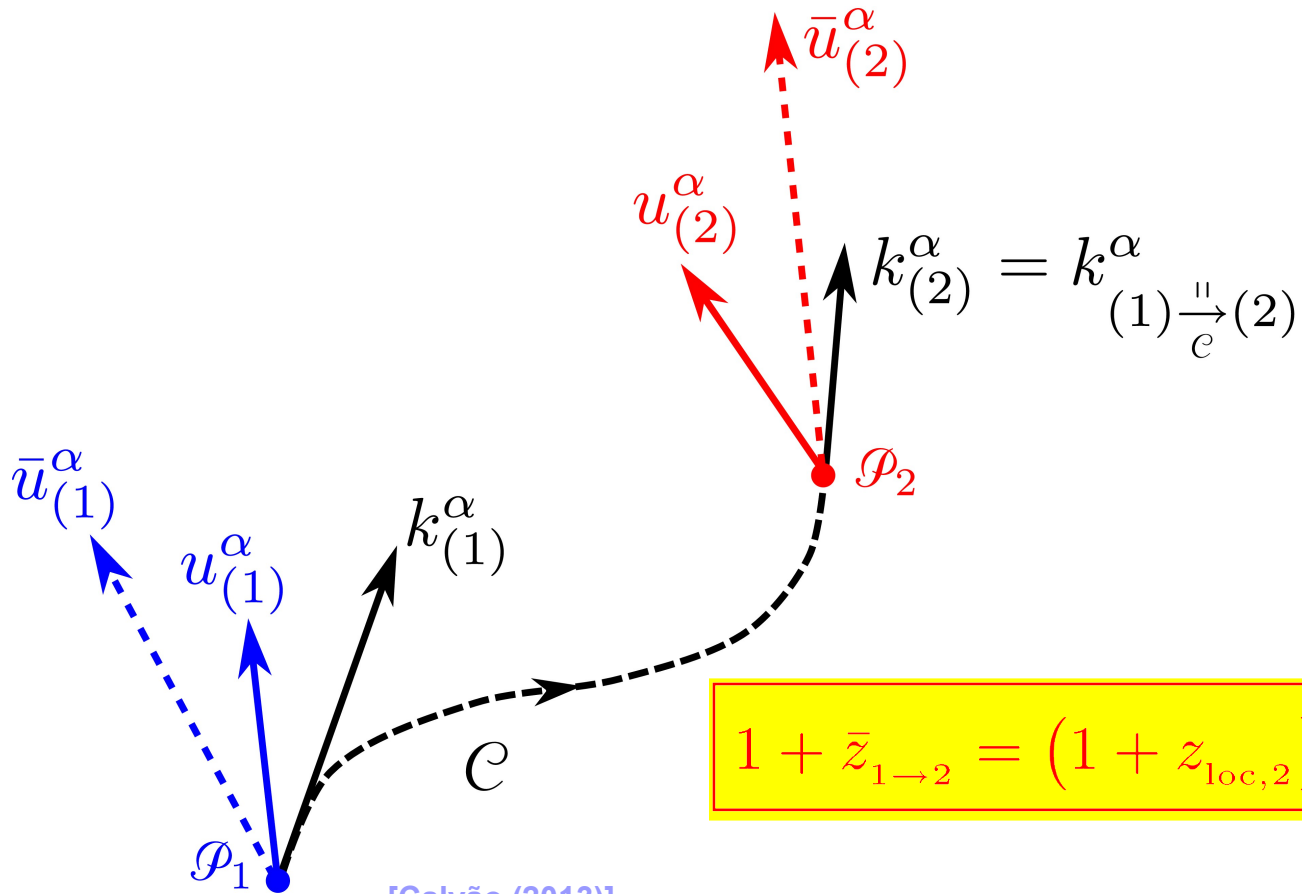


[Calvão (2013)]

II.3. REDSHIFT TRANSFORMATION



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$$1 + \bar{z}_{1 \rightarrow 2} = (1 + z_{\text{loc},2}) (1 + z_{1 \rightarrow 2}) (1 + z_{\text{loc},1})^{-1}$$

$$1 + \bar{z}_{1 \rightarrow 2} := \bar{\lambda}_2 / \bar{\lambda}_1 \qquad 1 + z_{1 \rightarrow 2} := \lambda_2 / \lambda_1$$

$$1 + z_{\text{loc},i} := \bar{\lambda}_i / \lambda_i$$



III. FLUX

III.1. DEFINITION

$$T^{\alpha\beta} = \rho u^\alpha u^\beta - Ph^{\alpha\beta} + 2q^{(\alpha} u^{\beta)} + \pi^{\alpha\beta}$$

$$f := ||q^\alpha||$$



geometrical optics (eikonal) approximation

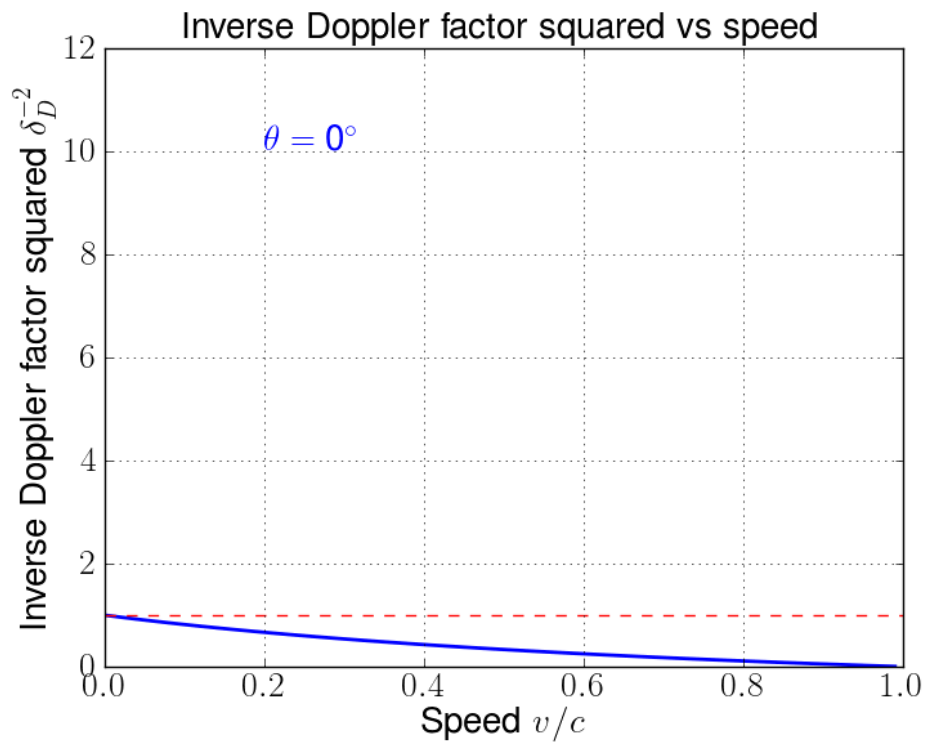
$$T^{\alpha\beta} = B^2 k^\alpha k^\beta \quad (k_\alpha := \partial\varphi/\partial x^\alpha)$$



$$f = \rho$$

III.2. FLUX TRANSFORMATION

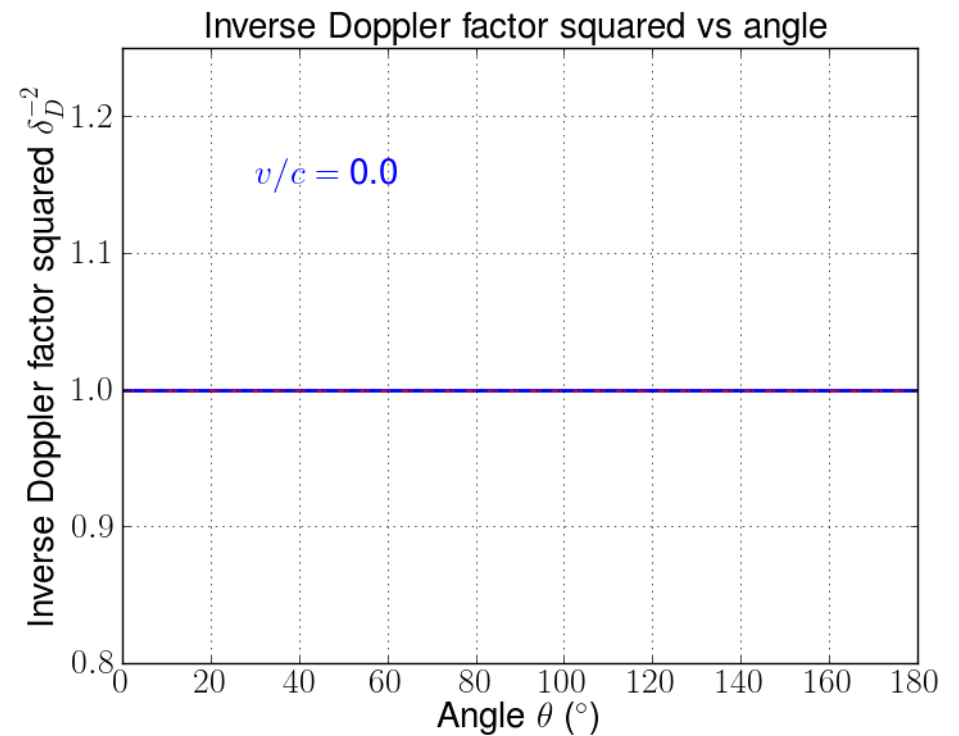
$$\bar{f} = f \delta_D^{-2}$$



[Calvão (2013)]

III.2. FLUX TRANSFORMATION

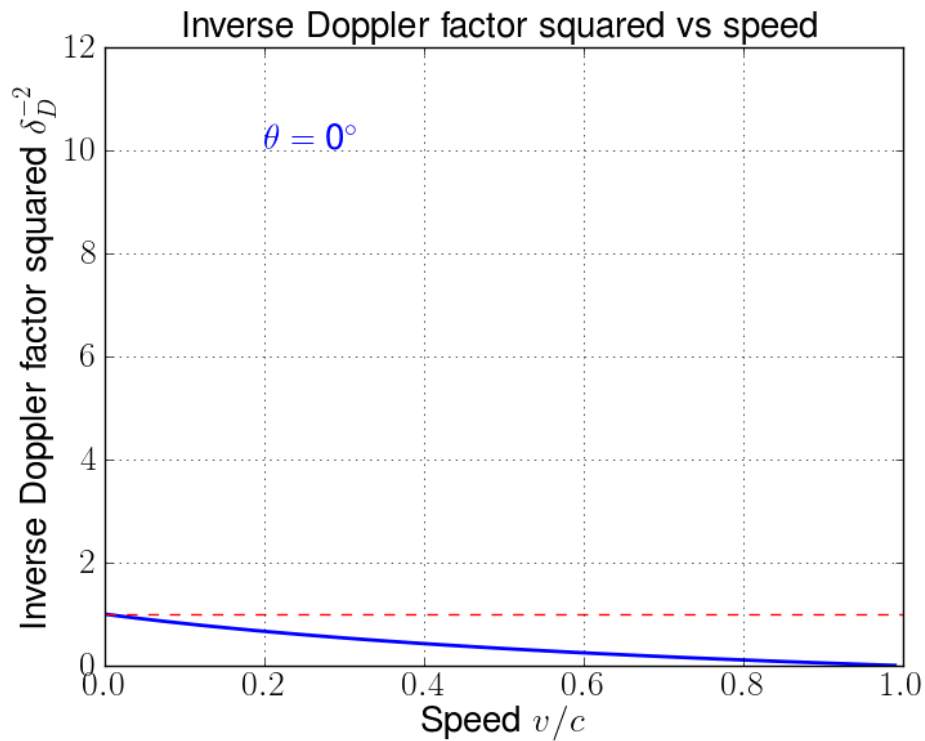
$$\bar{f} = f \delta_D^{-2}$$



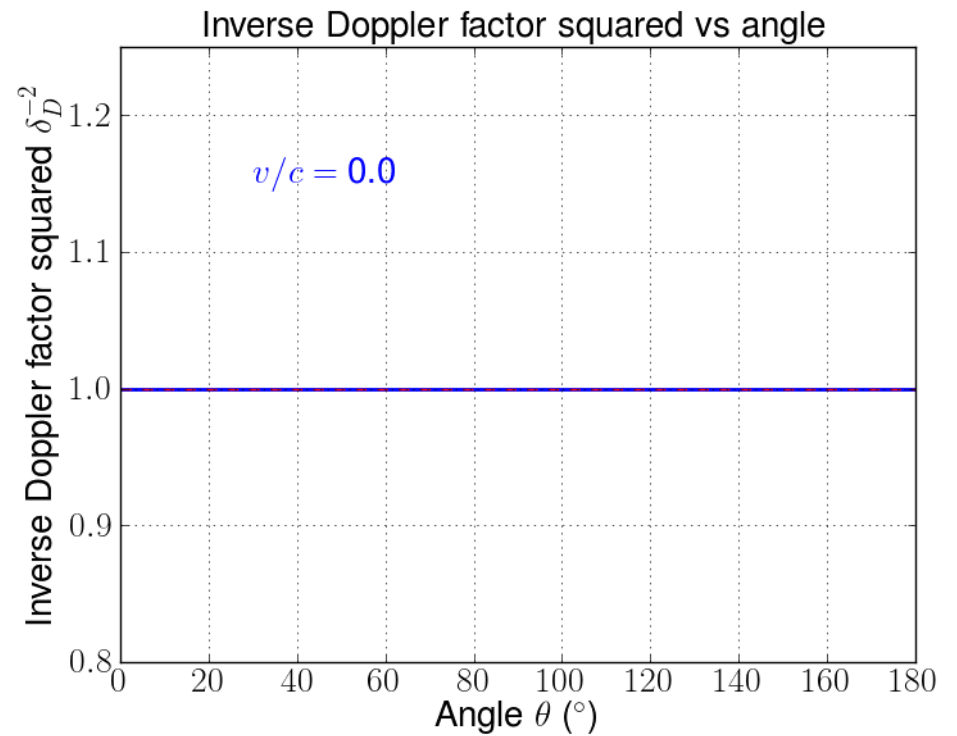
[Calvão (2013)]

III.2. FLUX TRANSFORMATION

$$\bar{f} = f \delta_D^{-2}$$



[Calvão (2013)]



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IV. LUMINOSITIES

IV.1. DEFINITIONS

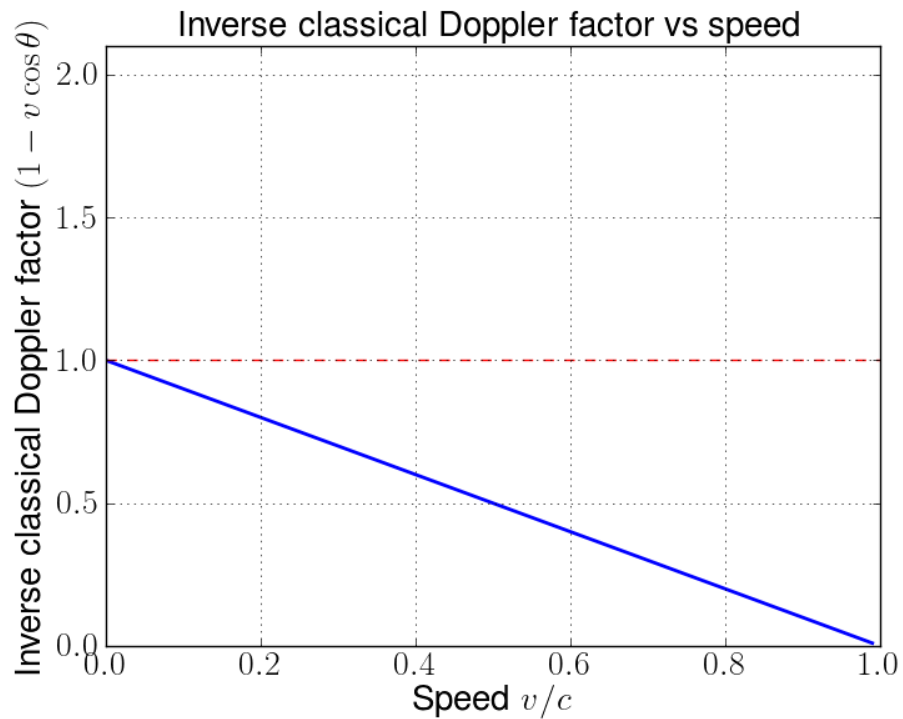
[Rybicki (1979)]

$L_{\Omega}(t, \hat{\Omega})$: angular distribution of emitted (source) power,
or simply angular emitted power

$\mathcal{L}_{\Omega}(t, \hat{\Omega})$: angular distribution of received (detector) power,
or simply angular received power

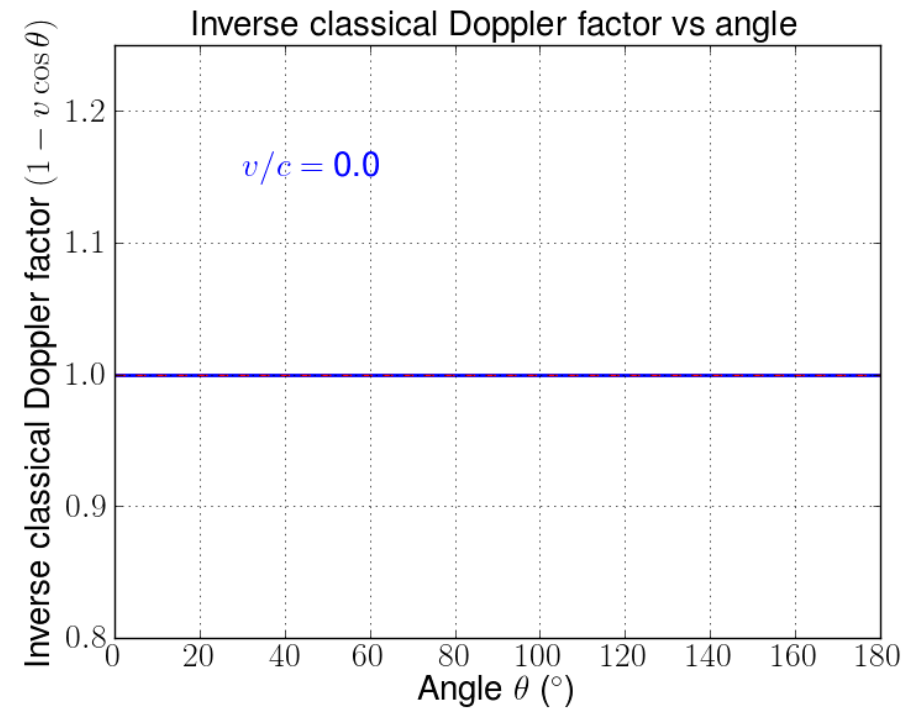


$$L_{\Omega} = \mathcal{L}_{\Omega}(1 - v \cos \theta)$$



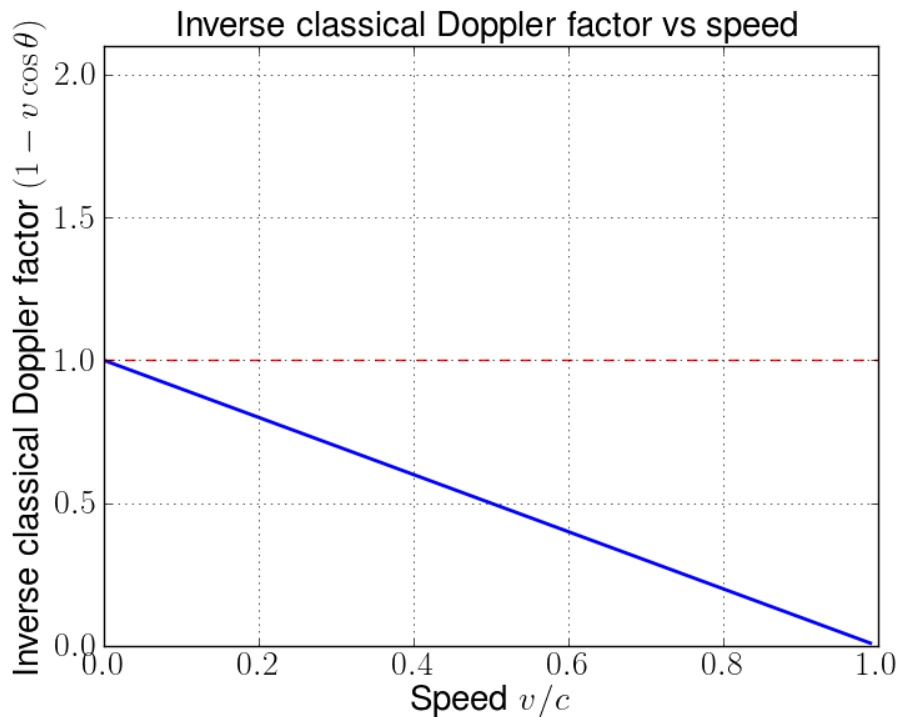
[Calvão (2013)]

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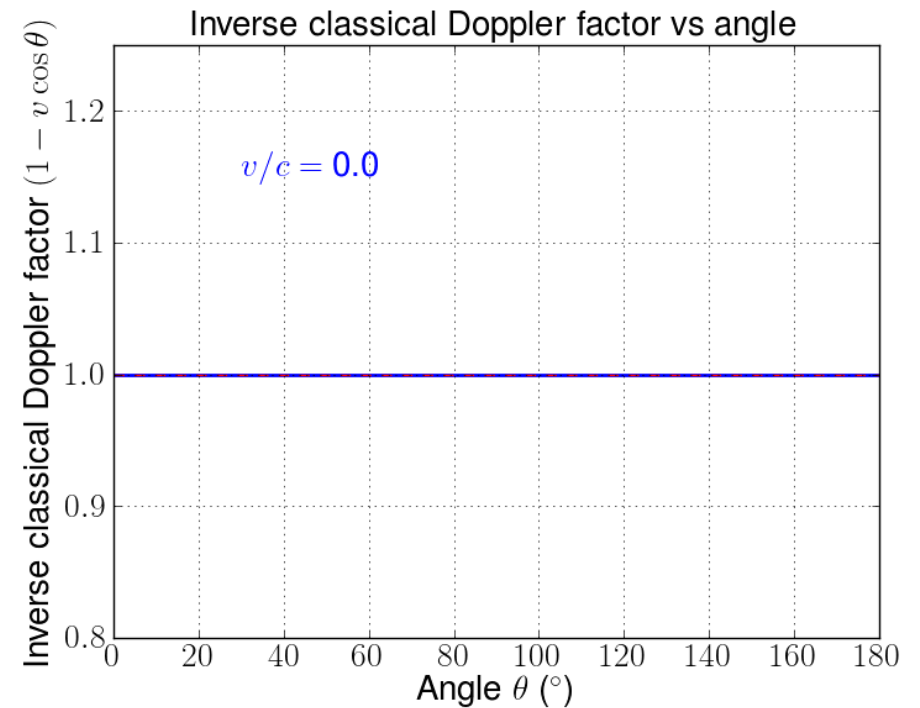


[Calvão (2013)]

$$L_{\Omega} = \mathcal{L}_{\Omega}(1 - v \cos \theta)$$



[Calvão (2013)]



[Calvão (2013)]



$L(t)$: emitted (source) power

$\mathcal{L}(t)$: received (detector) power

$$L = \mathcal{L} - v \oint \mathcal{L}(\hat{\Omega}) \cos \theta d\Omega$$

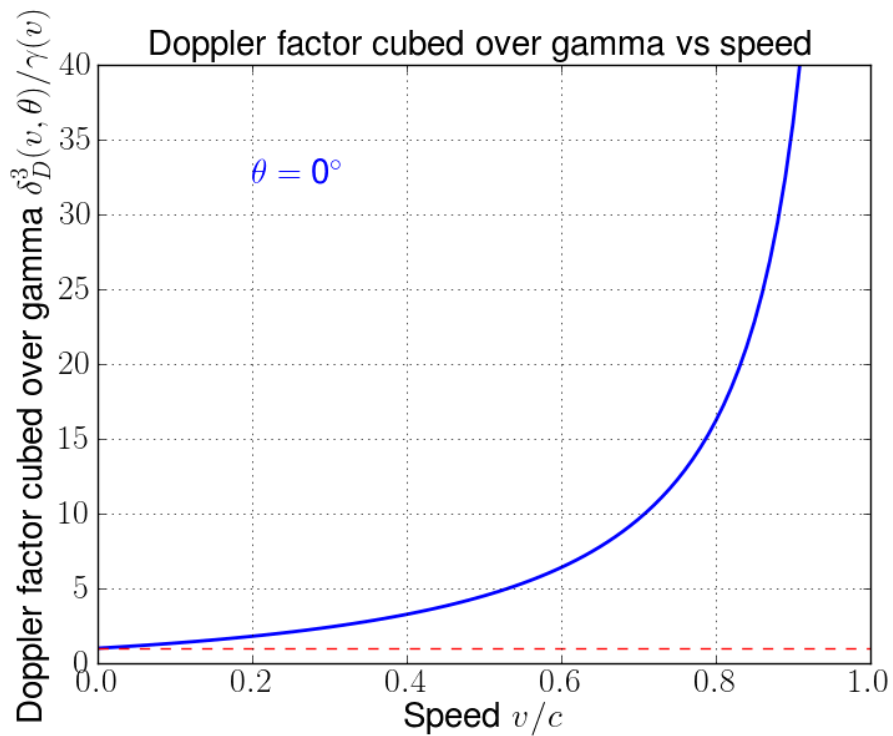
isotropic case:

$$L = \mathcal{L}$$



IV.1. TRANSFORMATION LAWS

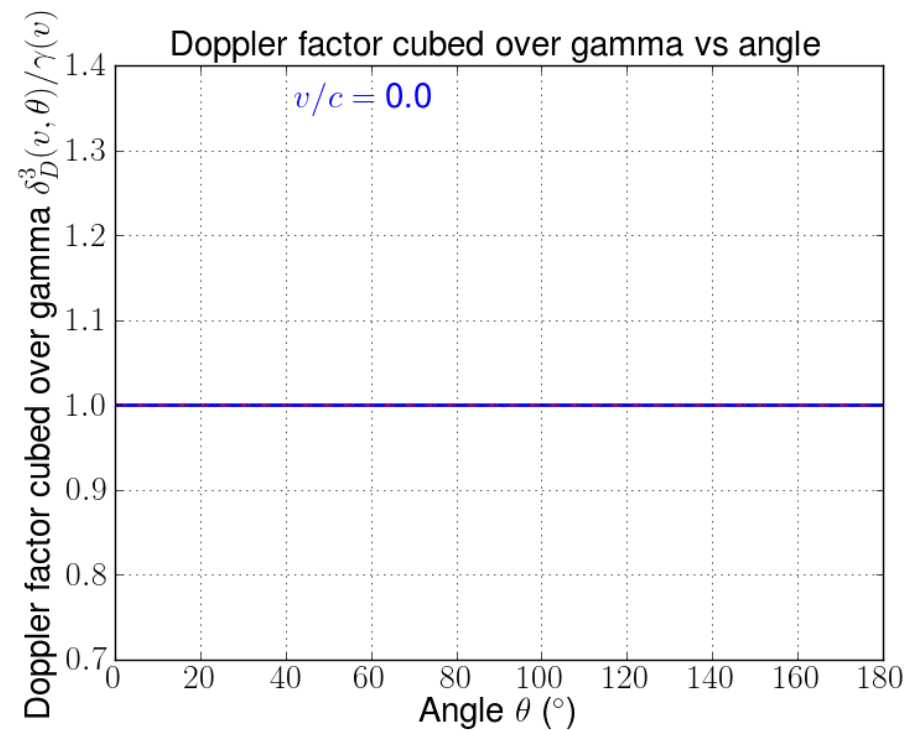
$$\bar{L}_\Omega = L_\Omega \gamma(v) \delta_D^3(v, \theta)$$



[Calvão (2013)]

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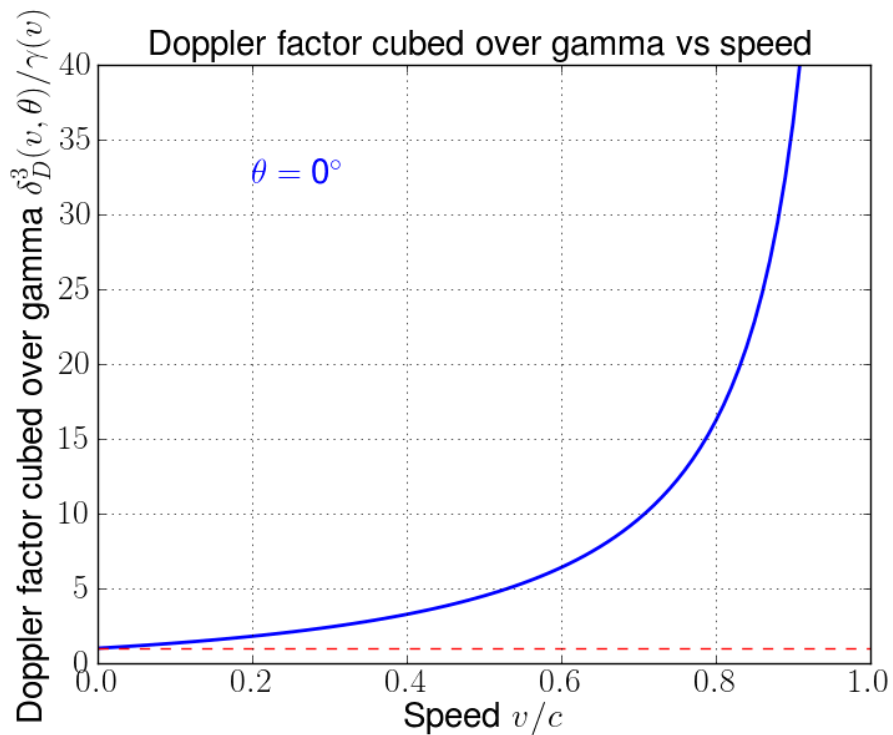
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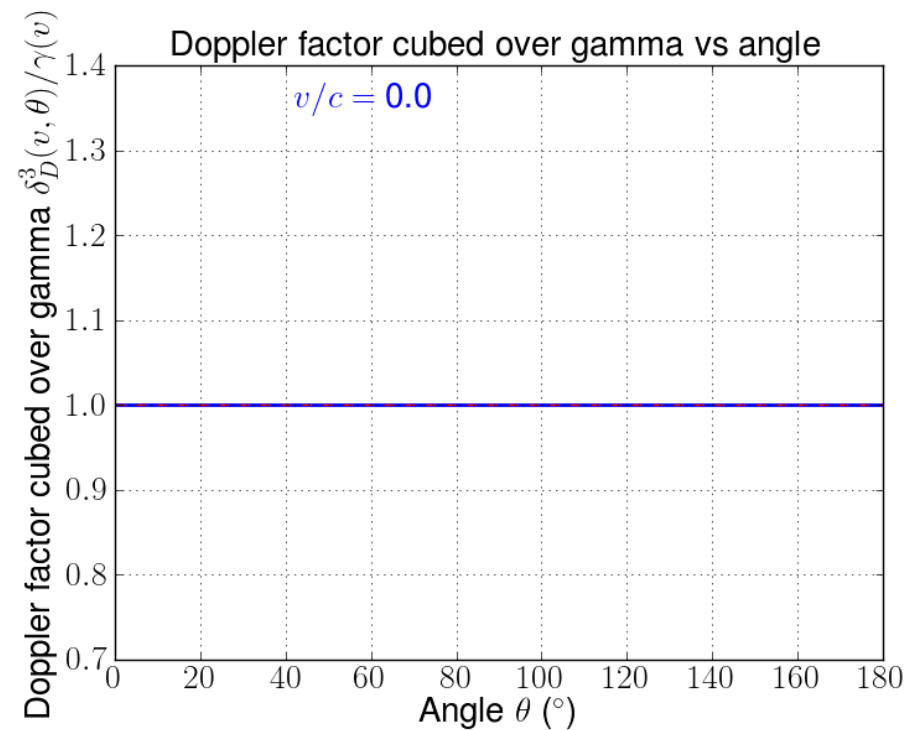
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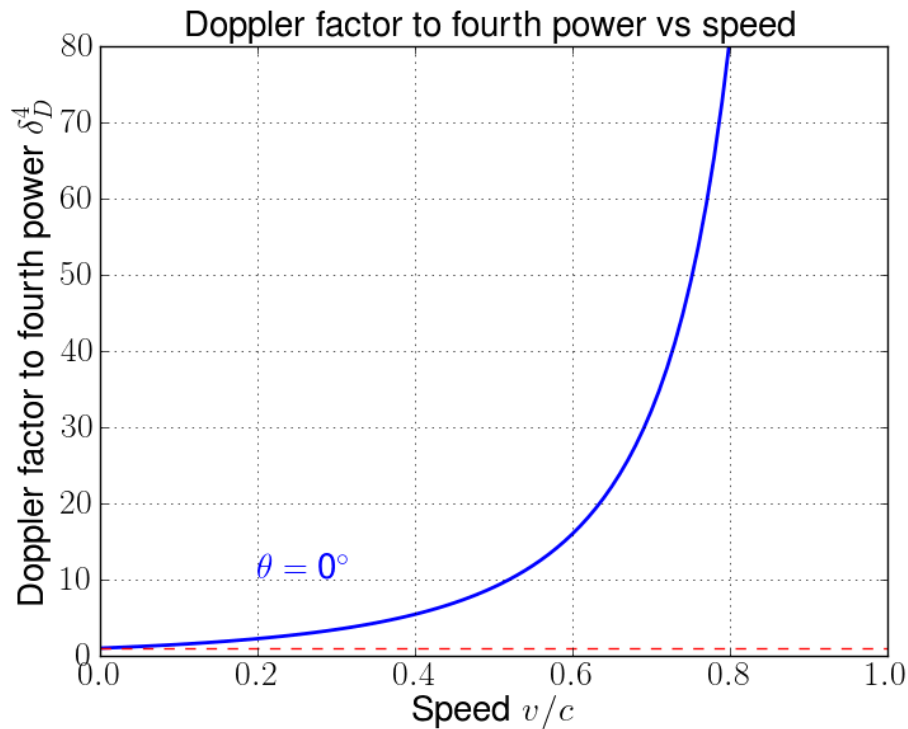
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$$\bar{\mathcal{L}}_{\Omega} = \mathcal{L}_{\Omega} \delta_D^4(v, \theta)$$

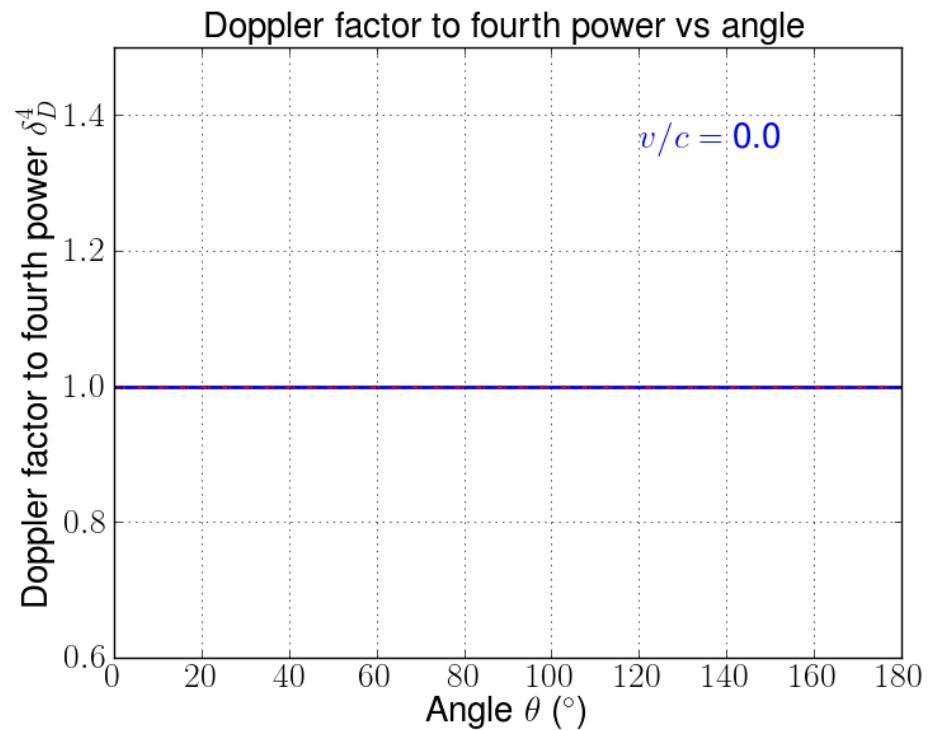


[Calvão (2013)]



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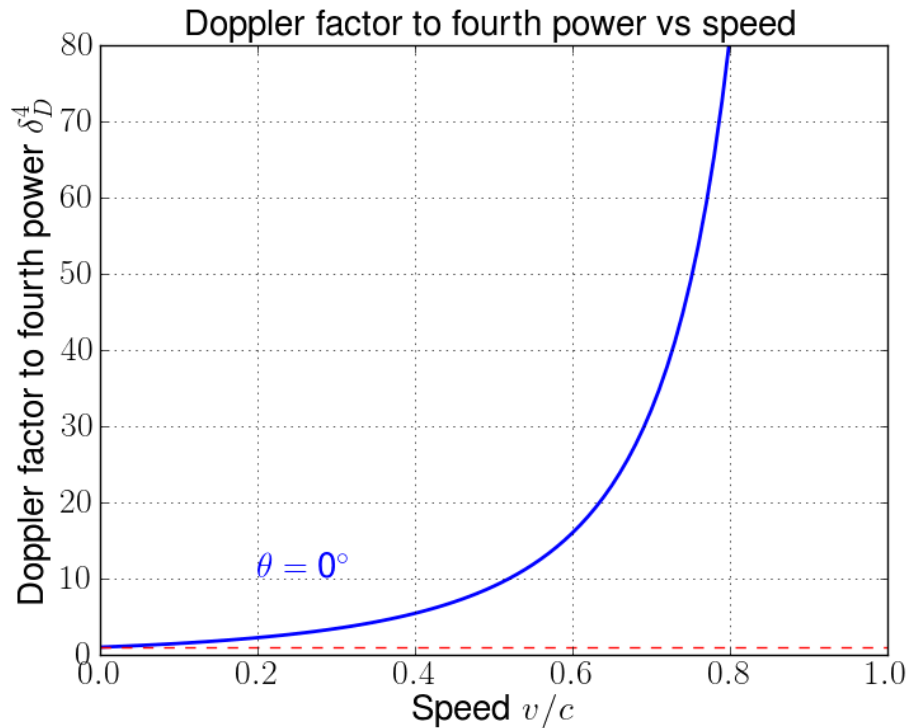
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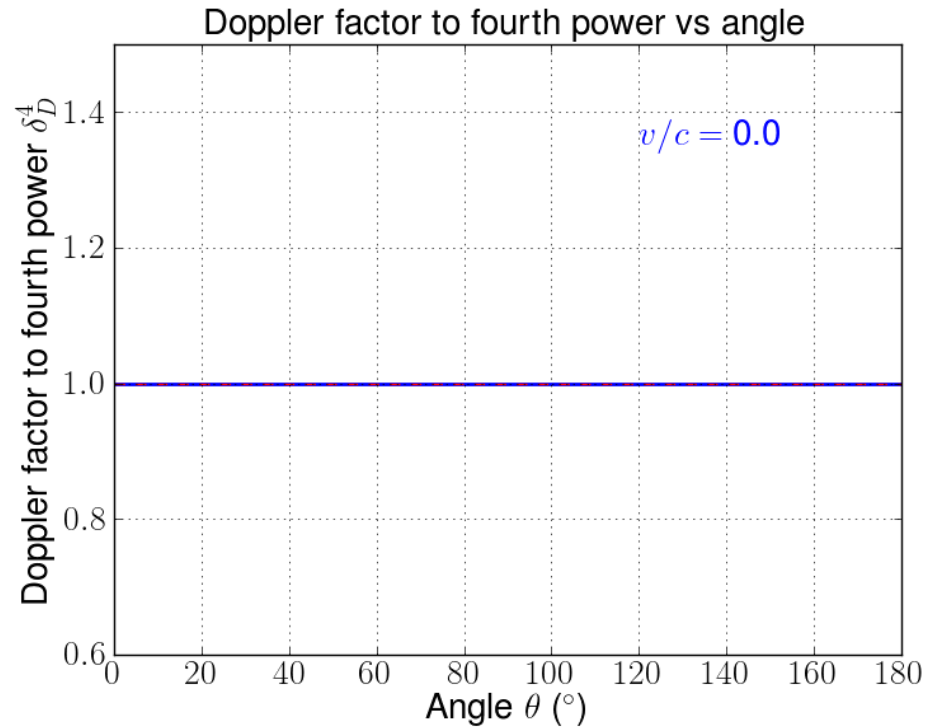
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$$\bar{\mathcal{L}}_{\Omega} = \mathcal{L}_{\Omega} \delta_D^4(v, \theta)$$



[Calvão (2013)]



[Calvão (2013)]



IV.1. TRANSFORMATION LAWS

$$\bar{L} = L + v \oint L_{\Omega}(\hat{\Omega}) \cos \theta d\Omega$$

$$\bar{\mathcal{L}} = \gamma^2 \left[\mathcal{L} + \oint \mathcal{L}_{\Omega} (2 + v \cos \theta) v \cos \theta d\Omega \right]$$

V. LUMINOSITY DISTANCES

V.1. DEFINITIONS

$$D_L := \sqrt{\frac{L}{4\pi f}}$$

$$D_{\mathcal{L}} := \sqrt{\frac{\mathcal{L}}{4\pi f}}$$

$$D_{L\Omega} := \sqrt{\frac{L\Omega}{f}}$$

$$D_{\mathcal{L}\Omega} := \sqrt{\frac{\mathcal{L}\Omega}{f}}$$

VI. CONCLUSION

MAIN RESULTS:

- We clarified the notions related to **redshift, flux, luminosities and luminosity distances** in a generic spacetime for arbitrary instantaneous observers, essentially without any approximation, not perturbatively!
- All Doppler effects can be interpreted as purely kinematic ones, arising from the use of non-parallelly transported instantaneous observers
- In particular, in any spacetime the redshift along a given null geodesic can always be made to vanish
- Flux is essentially the norm of the energy flux vector from the energy-momentum tensor and its transformation law is trivially obtained
- There are several consistent notions of (angular) luminosity and consequent luminosity distances
- Differences due to peculiar motions and/or anisotropies do arise and are not negligible at all!!